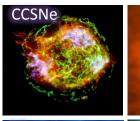
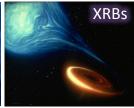
# **Astrophysics Working Group Summary**

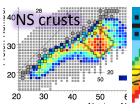
- Exciting, Broad, Robust Day One Science Program in Astrophysics
  - diverse topics include CCSNe, SNIa, Novae, XRBs, NS Mergers, Pop III stars
  - covering rp-, r-,  $\alpha$ p-,  $\nu$ p-, hot CNO, hot pp processes
- Broad Range of measurement types
  - capture, transfer,  $(\alpha,p)$ , decay, masses, charge exchange, total cross sections, TAS, LI fusion, HI collisions, HI fusion for EOS, fission ...
- Diverse set of beams
  - ReA3 beams <sup>30</sup>P, <sup>18</sup>F, <sup>22</sup>Mg, <sup>26</sup>Si, <sup>30</sup>S, <sup>56</sup>Ni, <sup>59</sup>Cu, <sup>34</sup>Ar, <sup>38</sup>K; <sup>8</sup>B, <sup>9</sup>C, <sup>13</sup>O, <sup>14</sup>O
  - ReA6 ReA12 beams <sup>30</sup>P, <sup>38</sup>K, <sup>59</sup>Cu, <sup>61</sup>Ga, <sup>65</sup>As; <sup>81</sup>Ni, <sup>76</sup>Cu, <sup>78</sup>Zn, <sup>80</sup>Ga, <sup>86,88</sup>As, <sup>131</sup>Cd, <sup>133 137</sup>Sn, <sup>137</sup>Te
  - Fast beams <sup>31</sup>F, <sup>125</sup>Tc, <sup>132</sup>Sn, <sup>19</sup>C, <sup>28</sup>Ne
  - Stopped beams near N=82 and N=126, Uranium
  - Stable beams (for commissioning) <sup>20</sup>Ne, <sup>15</sup>N, <sup>21</sup>Ne, Ti, Ni, Al, Mg
- Wide range of intensities
  - capture  $10^7$  pps, transfer & chargex  $10^5$  pps,  $\sigma$   $10^3$  pps, decay <<  $10^2$  pps
- Wide range of experimental stations in all FRIB halls
  - SECAR, JENSA, HRS, ISLA, HELIOS, GRETINA/GRETA, AT TPC, HR AT TPC ANASEN, S800, DECAY STATION, SUN, MUSIC
- Numerous facility requirements
  - suite of intense >10<sup>5</sup>-10<sup>6</sup> pps ReA3 beams with gas and solid stopper
  - beam purification methods; long-pulse beam time structure
  - small beam widths < 3 mm; use  $^{64}$ Zn primary beam for  $^{59}$ Cu ReA3 beam
  - DAQ: unified, digital, > 1000 channels
- Other requirements
  - H<sub>2</sub>, <sup>3</sup>He operations in JENSA; dispersion matched HRS with long flight path, precise 0.4 mm positions, <30 ps time resolution; GRETA with auxiliary detectors; fast detectors for high contamination ...

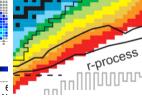




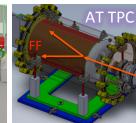


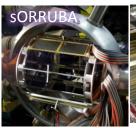


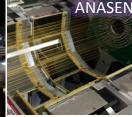


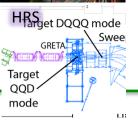












# **Astrophysics Working Group Summary**

#### General:

- aggressively develop intense ReA3 beams with a full suite of stopping and isotope collection
- many astrophysics experiments need >1e5 1e6 pps

### (p,γ) reactions direct measurements for rp-process with SECAR:

- >1e6 pps beams with gas and solid stopper technology. Examples: <sup>30</sup>P, <sup>38</sup>K, <sup>59</sup>Cu
- Suite of intense stable beams to commission SECAR
- Add <sup>64</sup>Zn primary beam to list

## • $(\alpha,p),(p,\alpha)$ reactions for rp/ $\alpha$ p/ $\nu$ p process: ANASEN, JENSA

- >1e6 pps of <sup>18</sup>F, <sup>22</sup>Mg, <sup>26</sup>Si, <sup>56</sup>Ni, <sup>59</sup>Cu reaccelerated beams
- development of techniques to deal with contamination purification, fast detection systems
- lower intensity (< 1e4 pps) ( $\alpha$ ,p) studies with MUSIC,  $^8B$ , $^8C$ ,  $^{14}O$ ,  $^{18}Ne$

# (d,p), (d,n), (<sup>3</sup>He,d) transfer for rp-process

- >1e5 pps beams, many beams
- 3He operation of JENSA
- decay studies: r-process around N=82 first, mostly Uranium beams
- Masses with TOF
  - HRS, dispersion matched optics mode, long flight path, precise tracking 0.4 mm, time detection with <30 ps resolution

## (p,y) resonance studies for rp-process with GRETA

- auxiliary detectors for neutron and charged particle detection needed for GRETA
- may need dedicated systems, DAQ integration
- Fission for r-process: Actinide fast beams
- EOS probes through heavy ion collision, fast <sup>132</sup>Sn beam
  - modify AT-TPC for intense fast beams, but could be done with system today
  - AT-TPC in fast beam HRS area (move magnet, or use different magnet)
- TAS for r-process (n,γ rates and beta decay)
- 12C + 19C and 20Ne + 28Ne fusion for neutron star crusts
  - <500 pps but exotic short-lived neutron rich beams issues of isobars from decay?</p>
- weak interaction strength with charge exchange for supernovae and neutron stars
  - Need HRS + TPC